Multifractal Dimension Spectra in Polymer Physics

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We study the multifractal properties of diffusion in the presence of an absorbing polymer and report the numerical values of the multifractal dimension spectra for the case of an absorbing self avoiding walk or random walk.

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1. Introduction

The notion of non-integer dimension although developed by mathematicians in late 19th - early 20th centuries has found its wide application in physics only in the last few decades. This application was mainly initiated by the pioneering work of B. Mandelbrot who attracted attention to the fact that "fractal" geometry provides an appropriate framework for the description of a whole range of complex structures that are formed from smaller subunits [1]. Whereas such structures are characterized by appropriate fractal dimensions, their growth and spatial correlations [2] are described by a (non-trivial) spectrum of multifractal (MF) dimensions [3, 4].

In this paper, we study the properties of diffusion phenomena in the presence of an absorbing polymer. This provides another example of a multifractal phenomenon in condensed matter physics. To this end we use the model proposed by Cates and Witten [5] and derive the MF spectrum in the frames of a field theoretical formalism using the renormalization group [6, 7] method. The MF spectrum is related to the spectrum of exponents governing scaling properties of co-polymer stars [8]. We calculate this spectrum to the third order of perturbation theory and report the numerical values of the quantities that characterize the MF behaviour.

2. The Model and the Multifractal Measure

Long flexible polymer chains [9] provide a classical example of fractal objects (see fig. 1a). Of particular interest is the manner in which such a chain absorbs diffusing

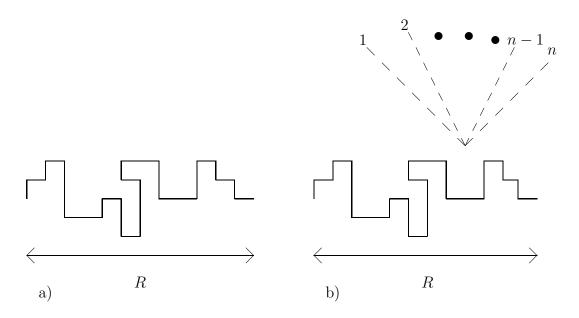


Figure 1. a) A SAW of N steps each of size a in d-dimensional space. In the asymptotic limit $N \to \infty$ the mean square distance between the extremities is related to the 'mass' N by $N \simeq (R/a)^{1/\nu}$ with $\nu(d=2)=3/4$, $\nu(d=3)=0.588$. This defines the fractal dimension of a SAW as D=4/3, D=1.70 in two and three space dimensions. b) A star of n RW (trajectories of diffusing particles) that reach the same point close to the absorbing polymer. The averaged nth moment of concentration of diffusing particles ρ scales as $<\rho^n>\sim (R/a)^{-\lambda_n}$. The set of scaling dimensions λ_n depends on n in a non-trivial way, leading to the non-trivial behaviour of the flux moments of diffusing particles on the surface of fractal (polymer) absorber.

particles. The flux of particles on the surface of the absorbing chain is defined by the density of incoming particles close to the absorber (see fig. 1b). Taken $\rho(r)$ to be the probability density of incoming particles the problem is to solve the steady-state diffusion equation (Laplace equation)

$$\nabla^2 \rho(r) = 0 \tag{1}$$

with boundary conditions for a given absorbing polymer

$$\rho(r) = 0$$
 on the surface of the absorber
$$\rho(r) = \rho_{\infty} = const for |r| \to \infty. (2)$$

One motivation for the study of this problem is to gain insight to the description of diffusion limited aggregation (DLA) [10]. The last phenomenon is much more

complicated because of the fact that for DLA the boundary conditions are given on the surface of growing aggregate itself. The process described by equations (1), (2) may be rather considered as diffusion limited catalysis, when particles of one type interact with a *prescribed* fractal (a catalysing polymer) and transform into the other type [12].

Let the absorbing polymer of size R be chosen from the well defined ensemble of SAW. Then we are interested in the moments of the field $\rho(r)$ close to the surface of the absorber (see fig. 1b). For distances R >> r > a (a being a cut off) the averaged moments $< \rho(r)^n >$ are expected to scale as [5]

$$<\rho(r)^n>\sim (R/r)^{-\lambda(n)},$$
 (3)

where $\langle ... \rangle$ denotes the average over the ensemble of polymers. Let us take that the flux $\phi(x)$ onto any randomly chosen point x of the absorber is proportional to the field $\rho(a)$ at a point that is as close as a cut off length a from the absorber. Then, for the averaged moments of the flux one finds:

$$<\phi^n>\sim (R/a)^{-\lambda(n)}.$$
 (4)

Following [5] we associate with the flux of the Laplacian field a "harmonic measure" $\mu(x)$ defined on the set of absorbing sites x:

$$\mu(x) = \phi(x) / \sum \phi(x), \tag{5}$$

where the sum in (5) spans all sites x of the absorber. In particular, the mass M of the absorber is given by $\sum_{x} 1 = M$ and (5) may be rewritten as

$$\mu(x) = \phi(x)/(M < \phi >), \tag{6}$$

where $\langle \ldots \rangle$ denotes the site average $\langle \ldots \rangle = \sum (\ldots)/M$ [11].

It is standard [3] to describe the properties of the harmonic measure by a set of exponents D(n) given by:

$$\sum \mu(x)^n = (R/a)^{(1-n)D(n)}. (7)$$

Note, that for n=0 D(0) is the fractal dimension of the absorber. Multifractal scaling is found when the spectrum D(n) is non-trivial: $D(n) \neq D(0)$, or correspondingly (c.f. eq. (4)) $\lambda(n) \neq n\lambda(1)$. Note as well that from (4), (7) the exponents $\lambda(n)$ and D(n) are related:

$$n\lambda(1) - \lambda(n) = (1-n)[D(n) - D(0)].$$
 (8)

In many cases the description of a multifractal measure is performed in terms of the Hölder exponent α and the spectral function $f(\alpha)$ that are defined by a Legendre transform of D(n) [4]:

$$\alpha(n) = \frac{d}{dn}[(n-1)D(n)], \tag{9}$$

$$f(\alpha) = n\alpha(n) + (1-n)D(n). \tag{10}$$

The function $f(\alpha)$ differs for different fractal measures but it possesses certain 'universal' properties. In particular, its maximum $\max_{\alpha} f(\alpha)$ gives the fractal dimension of the absorber and f is convex $f''(\alpha) < 0$.

3. The Multifractal Spectrum Exponents and the Spectral Function

The central idea that allows us to calculate the properties of the solutions of equation (1) with the boundary conditions (2) is that a path integral representation both for the field $\rho(r)$ and for the absorber (being a RW or a SAW) is possible 5, 12. In terms of the path integral solution of the Laplace equation one finds that $\rho(r)$ at point r near absorber is proportional to the number of RW that end at point r and avoid the absorber. The nth power of this field is proportional to the nth power of the above mentioned number, i.e. it is defined by a partition function of a star polymer with n arms [13] (the latter is shown by dashed lines in fig. 1b). Furthermore, introducing the mutual avoidance conditions between the "n-arm star" and the 2-arm polymer (representing the absorber) one has to calculate the partition function of a co-polymer star consisting of chains of two different species that avoid each other. These correspond to the trajectories of diffusing particles (being RW) and the absorbing polymer (which for the purpose of present study is chosen as RW or SAW). Making use of the previously developed [8] theory of copolymer stars and networks and mapping the model of co-polymer stars to the appropriate field theory [14, 15] one may relate [17] the MF spectrum exponents (4) to the exponents that define the scaling properties of co-polymer stars. In particular, for a co-polymer star consisting of n_1 chains of species 1 and n_2 chains of species 2 the number of configurations Z_* scales like [8]:

$$Z_* \sim (R/a)^{\eta_{n_1 n_2} - n_1 \eta_{20} - n_2 \eta_{02}}.$$
 (11)

Here, the $\eta_{n_1n_2}$ represent the co-polymer star exponents. The latter have been calculated using a field theoretical renormalization group approach [6, 7] and are known in the third order of perturbation theory in $\varepsilon = 4 - d$ and pseudo- ε [18] expansions [8, 16]. By means of a short-chain expansion [19] the set of exponents $\eta_{n_1n_2}$ can be related to the exponents $\lambda(n)$ (4), that govern the scaling of the *n*th moment of the flux onto an absorbing linear chain. Considering the absorber to be either a RW or a SAW we define the exponents:

$$\lambda_{RW}(n)^G = -\eta_{2n}^G,\tag{12}$$

$$\lambda_{SAW}(n)^U = -\eta_{2n}^U + \eta_{20},\tag{13}$$

here, we keep notations of [8, 16] for the expressions for exponents η defined in different fixed points of the renormalization group transformation. Based on these expressions we get:

$$\lambda_{RW}(\varepsilon) = \varepsilon n - \frac{n(n-1)\varepsilon^2}{4} + \frac{n(n-1)(-1+n+3\zeta(3))\varepsilon^3}{8}, \quad (14)$$

$$\lambda_{SAW}(\varepsilon) = \frac{3\varepsilon n}{4} + \frac{n(7n-18)\varepsilon^2}{128} + \frac{n(-149 - 42n + 108n^2 - 276\zeta(3) + 540\zeta(3)n)\varepsilon^3}{2048}.$$
 (15)

Table 1. Exponents $\lambda_{RW}(n)$ and $\lambda_{SAW}(n)$ obtained in ε and in pseudo- ε expansion $(\lambda(\tau))$ techniques.

n	$\lambda_{RW}(arepsilon)$	$\lambda_{RW}(au)$	$\lambda_{SAW}(\varepsilon)$	$\lambda_{SAW}(au)$
1	0.99	0.99	0.71	0.71
2	1.77	1.81	1.31	1.33
3	2.45	2.53	1.86	1.92
4	3.01	3.17	2.34	2.44
5	3.51	3.75	2.78	2.94
6	3.95	4.28	3.19	3.41

Here, $\varepsilon = 4 - d$, d is the space dimension, and $\zeta(3) \simeq 1.202$ is the Riemann zeta function.

Using the perturbation expansions for the λ exponents and the relations for $\lambda(n)$ and the spectral function some algebra leads to the corresponding expansions for α_n and $f(\alpha_n)$:

$$\alpha_{RW}(\varepsilon) = 2 + \frac{(-2n+1)\varepsilon^2}{4} + \frac{(-4n-3\zeta(3)+3n^2+6\zeta(3)n+1)\varepsilon^3}{8}, \quad (16)$$

$$f_{RW}(\varepsilon) = 2 - \frac{\varepsilon^2 n^2}{4} + \frac{(2n^3 + 3\zeta(3)n^2 - 2n^2)\varepsilon^3}{8},$$
 (17)

$$\alpha_{SAW}(\varepsilon) = 2 - \frac{\varepsilon}{4} + \frac{(7 - 36n)\varepsilon^2}{128} + \frac{(-149 - 276\zeta(3) - 84n + 324n^2 + 1080\zeta(3)n)\varepsilon^3}{2048},$$
(18)

$$f_{SAW}(\varepsilon) = 2 - \frac{\varepsilon}{4} - \frac{(18n^2 + 11)\varepsilon^2}{128} + \frac{(216n^3 + 540\zeta(3)n^2 - 42n^2 - 83 + 264\zeta(3))\varepsilon^3}{2048}.$$
 (19)

The exponents $\lambda(n)$ as well as the Hölder exponents α and the spectral functions $f(\alpha)$ in terms of the pseudo- ε expansion [18] are given elsewhere [17, 20].

As is well known, the series of type (14)-(19), as they occur in field theory appear to be of asymptotic nature with zero radius of convergence. However, knowledge of the asymptotic behavior of the series as derived from the renormalization group theory allows us to evaluate these asymptotic series (see e.g. [7]). To this end several procedures are available depending on the additional information known for

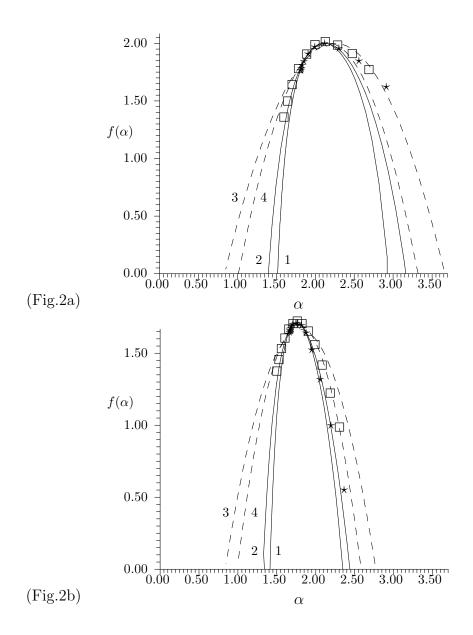


Figure 2. The spectral function $f(\alpha)$ for absorption on (a) RW and (b) SAW. Solid curves: 1 - [2/1] Padé approximant for ε^3 results, 2- [2/1] Padé approximant for pseudo- ε^3 results; dashed curves: 3 - ε^2 results without resummation, 4 - pseudo- ε^2 results without resummation; stars - resummed ε^3 results; boxes - resummed pseudo- ε^3 results.

the series to be resummed. We use the Borel summation technique improved by a conformal mapping procedure that has served as a powerful tool in field theoretic calculations (see [21] for example).

Table 1 contains the results for the exponents $\lambda_{RW}(n)$ and $\lambda_{SAW}(n)$ obtained in ε and in pseudo- ε expansion techniques at d=3 from the corresponding values of exponents η_{f_1,f_2} [8] with the application of this resummation procedure.

4. Conclusions

We have calculated the spectral function that describes the scaling of the moments of measure defined by diffusion near an absorbing polymer. These moments were calculated as averages over all configurations of the absorber instead of performing a site average. Thus the interpretation of $f(\alpha)$ itself does not directly correspond to the standard picture, where $f(\alpha)$ is interpreted as the fractal dimension of a particular subset of the measure [4]. Nevertheless we derive $f(\alpha)$ in the same way from the spectrum of scaling exponents of the moments of a measure. As we will see below in the general properties of $f(\alpha)$ also hold for the present definition.

Our numerical results for the spectral function for space dimension d=3 are presented in Figs. 2a,2b. They were obtained from the series for α_n and $f(\alpha_n)$ as functions of n. We show the results of the resummation procedure described above applied to the series in ε - and pseudo- ε - expansion. For comparison we also show the curve for direct summation of the ε and pseudo- ε series to the 2nd order (here we recover the second-order results of [5]). In addition we have performed an analytical continuation of our series in the form of a [2/1] Padé approximant for the third order series. It is obvious that the direct summation of the last series fails to converge and gives reasonable values for α_n , $f(\alpha_n)$ only for small values of n, i.e. near the maximum of $f(\alpha)$ at n=0. The symmetry of the Padé approximant holds only in the region shown and may be an artifact of the method. On the left wing, where it coincides with the resummed results the Padé approximant provides a continuation that is compatible with the estimation for the minimal α value $\alpha_{\min} = d - 2$. The Padé result is $\alpha_{\min}(\varepsilon) = 1.333$, $\alpha_{\min}(\tau) = 1.017$ for the RW absorber and $\alpha_{\min}(\varepsilon) = 1.250$, $\alpha_{\min}(\tau) = 1.013$ for the SAW absorber, which is calculated here only from 3rd order perturbation theory.

As can be extrapolated from the Padé approximant and as was shown also on the basis approximations for higher moments n [5], $f(\alpha)$ as it is defined here, will become negative near α_{\min} and α_{\max} . For this reason the identification of $f(\alpha)$ as the fractal dimension of some identifiable subset is not possible here. Also the extrapolation of the resummed data seems to indicate such a behavior. Note, however, that the perturbative approach, even in combination with resummation and analytical continuation is still not capable to give reliable results for high values of the expansion parameters. In particular this method is only good near the maximum of $f(\alpha)$ [22].

Though the results obtained for α_n and $f(\alpha_n)$ for a specific value of n differ in both approaches, the same curve $f(\alpha)$ is described with better coincidence for the

left wing of the curves, corresponding to positive n. The resummation techniques we apply have proven to be a powerful tool already in the field theoretical approach to critical phenomena and have lead to high precision values for critical exponents. We believe that the application of these methods to calculations of MF phenomena also allows to improve the reliability and comparability of these results.

A generalization of our present approach to the study of spectral functions that describe the absorption at the core of an absorbing polymer star with a given number of arms will be published elsewhere.

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